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Today, paleomagnetism is a mature science. Our methods, analyses, and results are more sophisticated than they were 30 years ago. Therefore, we feel it is appropriate to revisit the Van der Voo (1990) criteria in light of those developments. We hope to honor the intention of the original paper by keeping the criteria simple and easy to evaluate while also acknowledging the advances in science. This paper aims to update the criteria and modernize the process. We base our changes on advances in paleomagnetism and geochronology with a faithful adherence to the simplicity of the original publication. We offer the "Reliability" or "R" index as the next generation of the Van der Voo "Quality" or "Q" index. The new R-criteria evaluate seven different information items for each paleomagnetic pole including age, statistical requirements, identification of magnetic carriers, field tests, structural integrity, presence of reversals and an evaluation for possible remagnetization.

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The Magnificent Seven: A Proposal for Modest Revision of the Van der Voo (1990) Quality Index

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Abstract

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Today, paleomagnetism is a mature science. Our methods, analyses, and results are more sophisticated than they were 30 years ago. Therefore, we feel it is appropriate to revisit the Van der Voo (1990) criteria in light of those developments. We hope to honor the intention of the original paper by keeping the criteria simple and easy to evaluate while also acknowledging the advances in science. This paper aims to update the criteria and modernize the process. We base our changes on advances in paleomagnetism and geochronology with a faithful adherence to the simplicity of the original publication. We offer the "Reliability" or "R" index as the next generation of the Van der Voo "Quality" or "Q" index. The new R-criteria evaluate seven different information items for each paleomagnetic pole including age, statistical requirements, identification of magnetic carriers, field tests, structural integrity, presence of reversals and an evaluation for possible remagnetization.

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1. Introduction

47 Paleomagnetic studies were a crucial element in verifying continental mobility and the
48 establishment of plate tectonics as the prevailing paradigm in Earth Sciences in the 1950's and
49 1960's (Cox and Doell, 1960; Opdyke, 1995 and references therein). Nowadays, paleomagnetic
50 inquiry yields a quantitative assessment of plate motion and true polar wander over the bulk of
51 geologic time, and is the primary evidence used in quantitative continental reconstructions.
52 Furthermore, paleomagnetic data forms the basis for evaluating the evolution of the geodynamo.
53 As the number of paleomagnetic results increased and techniques for isolating magnetic
54 components were refined, it became apparent that not all studies were equally reliable. A
55 number of 'filtering' techniques were proposed to assess the reliability of an individual result
56 (Irving et al., 1976; Briden and Duff, 1981; May and Butler, 1986; Piper, 1987; Pesonen et al.,
57 1989; Li et al., 1990; Buchan et al, 2000). Additional rejection/selection criteria were sometimes
58 used on an ad-hoc basis (see for example Westphal and Pfaff, 1986), but it was not until Van der
59 Voo's (1990) study that a schema for grading paleomagnetic poles was widely applied.

60 Three parallel developments made the Van der Voo (1990) criteria immediately relevant
61 to the paleomagnetic community. The first "Nordic Paleomagnetic workshop" took place in
62 Espoo, Finland, in 1986, and the second in Sweden in 1990 (Elming and Pesonen, 2009). These
63 workshops brought together experts in paleomagnetism from the U.S.A. and Europe to review
64 paleomagnetic poles and evaluate their quality and reliability. The first two workshops, though
65 limited in scope to critical reviews of paleomagnetic data from Baltica and Laurentia, provided
66 the framework for subsequent global expansion. The second important development to arise in
67 this time frame was the hypothesis of the Precambrian supercontinent Rodinia (McMenamin and
68 McMenamin, 1990; Dalziel, 1991; Moores, 1991; Hoffman, 1991). Precambrian paleomagnetic
69 studies were relevant to the Rodinia hypothesis, but the quality of the database was viewed with
70 skepticism (Van der Voo and Meert, 1991; Piper, 1987). Perhaps the most important impetus
71 for advancing a grading scheme of paleomagnetic data was the creation of the Global
72 Paleomagnetic Database (GPMDB, McElhinny and Lock, 1990; Lock and McElhinny, 1991).
73 Global compilations of paleomagnetic data were available prior to the GPMDB (for example
74 Irving, 1960; McElhinny, 1968; Khramov, 1971), but they were neither digital nor easily
75 searchable. McElhinny and Lock (1990) purposely avoided grading individual poles, but the
76 compilation provided an easily queried database for critically evaluating published
77 paleomagnetic data. Rob Van der Voo was involved in all three prongs of these overlapping

78 research foci and had the forethought to develop a scheme for evaluating data leading to a critical
79 appraisal of our scientific approach to paleomagnetic studies.

80 2. *The Magnificent Seven*

81 Van der Voo (1990) proposed seven criteria to evaluate individual paleomagnetic studies,
82 given in a simplified form in Table 1. The original criteria were fleshed out by Van der Voo
83 (1993) in a discussion of (largely Cambrian and younger) paleomagnetism and paleogeography.
84 The following review of those criteria is couched within the knowledge base of the early 1990s.

85 The first criterion is met when the age of the magnetization is the same as the age of the
86 rock to within a half-period (or $\pm 4\%$, whichever is larger) for the Phanerozoic and to within $\pm 4\%$
87 (or 40 Ma, whichever is smaller) for the Precambrian. These age limits were based on
88 comparing average rates of apparent polar wander with uncertainties surrounding mean poles of
89 Phanerozoic and Paleoproterozoic age.

90 The second criterion established a statistical norm for the precision of paleomagnetic data
91 based on the number of samples required ($N \geq 25$), the clustering parameter k for directional data
92 means (Dec, Inc) or K for the of virtual geomagnetic poles (VGP's), where $k(K) \geq 10$ and the
93 cone of 95% confidence about the mean direction α_{95} (or A_{95} for paleomagnetic poles) is $\leq 16^\circ$
94 (Fisher, 1953).

95 Adequate demagnetization techniques used to isolate mean vectors forms the basis for the
96 third criterion. This requires stepwise alternating field, thermal or chemical demagnetization
97 techniques that can separate multicomponent magnetizations through the use of principal
98 component or great-circle analyses (Zijderveld, 1967; Halls, 1976, 1978; Kirschvink, 1980).

99 Criterion number four is met when the study can constrain the age of magnetization via a
100 field test. One of the commonly used field tests is the fold test (McElhinny, 1964), wherein the
101 age of magnetization is confirmed to be older/younger (or coeval with) the deformational event
102 that resulted in tilting/folding of the rocks. The fold test (Graham, 1949; Figure 1) is most
103 useful when the age of folding is only slightly younger than the rocks. Statistical complications
104 of the fold test are discussed at length in McFadden and Jones (1981), McFadden (1990),
105 McFadden (1998), Watson and Enkin (1983), Tauxe and Watson (1994) and Enkin (2003).

106 **Fig. 1 Here**

107 The baked contact test (Everitt and Clegg, 1962) is relatively straightforward in concept,
108 though often problematic in the field. Figure 2a shows a theoretical field setting of a dyke

109 intrusion and the expected thermal imprint on the host rocks. A positive baked contact test
110 should ideally include all of the following (a) a stable high-unblocking temperature (T_{ub})
111 magnetization in the intrusive body; (b) a stable high T_{ub} magnetization in the baked zone; (c) a
112 stable, but lower T_{ub} magnetization in the ‘hybrid’ (or partially baked) zone that matches the
113 direction in the intrusion and a stable high T_{ub} that matches the stable host rock direction.
114 Ideally, the T_{ub} of the dyke component should decrease with increasing distance away from the
115 contact and; (d) a stable high T_{ub} magnetization in the host rock that is distinct from the intrusive
116 body (figure 3b). Schwarz (1977) proposed a baked contact profile test which requires a more
117 detailed sampling profile through the intrusive body, contact, and host rocks in order to
118 demonstrate the acquisition of a partial thermal remanent magnetization (pTRM) that decreases
119 with distance from the contact (McClelland-Brown, 1981; Hyodo and Dunlop, 1993; Buchan et al.
120 (1993).

121 **Fig. 2 here**

122 Graham (1949) introduced the conglomerate test (Figure 3), positing that clasts of
123 parental rocks should have their directions randomized during subsequent transport. A test for
124 randomness was suggested by Bruckshaw and Vincenz (1954) and later formalized and
125 quantified by Watson (1956a,b; see also Irving, 1964; Stephens, 1964; Gine, 1975; Diggle et al.,
126 1985; Shipunov et al., 1998; Heslop et al., 2018). A positive test requires that the clasts have
127 randomly oriented directions as compared to the underlying or adjacent lithology from which the
128 clasts were derived. The age of the conglomerate (and its clasts) along with its relation to
129 bounding units is a critical consideration for evaluating the significance of the test. The
130 conglomerate test is best defined when the conglomerate is interbedded with the unit being
131 investigated (intraformational conglomerate test; MacNiocaill, 2000; Meert et al. 2009;
132 Levashova et al., 2009). Ideally, the demagnetization behavior (e.g. unblocking temperature
133 and/or coercivity) and rock magnetic characteristics of the conglomeratic clasts should be
134 identical to their parent material (Buchan and Hodych, 1989; Meert et al., 1994).

135 **Fig. 3 Here**

136 The fifth Van der Voo (1990) quality criterion was developed to address the reality that
137 paleomagnetic results from orogenic belts or from non-stratified (e.g. plutonic and metamorphic)
138 rocks can be problematic. Carey (1958) realized that fold belts can experience simple rotations
139 about a vertical axis. In the absence of a stable reference frame (craton), vertical axis rotations

140 can still provide critical information regarding paleolatitude (Van der Voo and Channell, 1980).
141 Without additional information, most plutonic and metamorphic rocks lack a suitable reference
142 for paleohorizontal and therefore are less likely to provide useful information regarding the
143 paleoposition of that block. Dyke swarms that preserve widespread verticality of the intrusions,
144 especially among intersecting swarms of variable orientations, are a notable exception to this
145 rule. Layered intrusions might also preserve useful paleohorizontal datums, although significant
146 tilting may occur between establishment of the igneous layering (ca. 1000°C) and the acquisition
147 of magnetic remanence ($\leq 580^\circ\text{C}$ for $\text{Fe}_{3-x}\text{Ti}_x\text{O}_4$).

148 For his sixth criterion, Van der Voo (1990) made the observation that paleomagnetic
149 studies that carry a dual-polarity magnetization provide evidence that (a) secular variation is
150 likely to be time-averaged and (b) is more commonly observed in rocks that are otherwise known
151 to carry a primary magnetization, although remagnetized rocks do sometimes carry a dual-
152 polarity magnetization (see Johnson et al., 1984; Johnson and Van der Voo, 1989). The presence
153 of polarity changes in sequential stratigraphic order provides the most powerful evidence for a
154 primary magnetization. Van der Voo (1990, 1993) acknowledged that his use of reversals as a
155 reliability criterion did not require a positive statistical test since the tests available at the time
156 were known to be flawed (e.g. Cox and Doell, 1960; McElhinny, 1973) or had not yet been
157 sufficiently tested (McFadden and McElhinny, 1990). In addition, the statistical tests merely
158 demonstrate the presence or absence of an incompletely removed secondary component of single
159 polarity, which might bias the characteristic remanence component such that its means
160 significantly differ from the 180° ideal case; in some situations, magnetostratigraphy may retain
161 its utility and robustness even if the statistical test fails (e.g. Evans et al., 2000).

162 The final, seventh Van der Voo (1990) criterion suggests the rejection of any
163 paleomagnetic pole that resembles a younger pole (>period) from the same craton or block. The
164 logic is simple in that resemblance to a younger paleopole from the same tectonic block raises
165 suspicion of a remagnetization. Van der Voo (1990) understood the significance of research
166 showing that remagnetization was common and not necessarily restricted to the tectonically
167 active margins of cratonic regions (McCabe and Elmore, 1989). Van der Voo (1990) argued
168 that field tests are required to ameliorate any concerns about remagnetization when an older pole
169 resembled a much younger pole from the same block.

170

171 **Table 1. Here**172 3. *The “R” Reliability Index: A Modest Proposal*

173 We propose a modest revision to the Van der Voo (1990) “Quality” Index. Colloquially
174 known as the “Q” factor, the paper significantly impacted the paleomagnetic community.
175 Paleomagnetic studies and proposals were framed in such a way as to meet as many of the Q-
176 criteria as necessary. Publications touted results as “our data earn a Q-value of X”. This
177 represented a major step forward in our science, but also created a number of debates amongst
178 the scientists responsible for evaluating the paleomagnetic database as well as interpreting the
179 original intent of Van der Voo (1990). Our goal is to review/revise these criteria in light of
180 modern methods, equipment and understanding of the science. We choose the letter “R” for this
181 scheme as it sequentially follows “Q” and also because “Reliability” is an accurate descriptor.

182 *Reliability Criterion #1-Age of the Rocks constrained to within +/- 15 Ma and*
183 *magnetization is presumed to be the same age as the rocks.*

184 We propose stricter age constraints for meeting the requirements of R1. Geochronological
185 studies have advanced in the past 28 years, particularly with respect to dating of mafic igneous
186 bodies using baddeleyite and zircon (e.g., Kamo et al., 1989). Geochronologists are also more
187 skilled at recognizing interbedded ash flows in sedimentary rocks (see, for example, Compston et
188 al., 1992; Grotzinger et al., 1995; Rasmussen et al., 2002). New techniques for direct dating of
189 sedimentary rocks are still in their nascent stages but show promise of providing robust age
190 control (McNaughton et al., 1999; Rasmussen et al., 2004; Selby and Creaser, 2005; Zhang et al.,
191 2015; Aleinikoff et al., 2015).

192 The original Q-scheme required that the age of a Phanerozoic paleomagnetic pole should be
193 constrained to within half a period (or $\pm 4\%$ *whichever is larger*) and that Precambrian poles
194 should be dated to within either 4% or ± 40 Ma (*whichever is smaller*). The Phanerozoic age
195 limits were based on apparent polar wander rates for Wrangellia and North America in the
196 Phanerozoic. The calculated average APW rate of $\sim 0.32^\circ/\text{Ma}$ degrees (3.5 cm/yr) resulted in
197 overlapping mean poles in 25-Ma windows. Van der Voo (1993) concluded that higher
198 precision ages would not result in better APWP resolution. The Precambrian threshold was
199 similarly determined on the basis of observed angular uncertainties of $\pm 16^\circ$ per 80 Ma (Van der
200 Voo and Meert, 1991). This seemingly simple scheme nevertheless produces some rather odd
201 results. The Cretaceous Period spans about 80 Ma. Therefore, a Cretaceous-age pole with an

202 error of ± 40 Ma would be acceptable (and equivalent to the acceptable maximum error of a
203 Precambrian-age pole). In contrast, a pole from the Silurian (which spans ~ 25 Ma) would
204 demand an error of less than ± 13 Ma in order to meet Q1. Because of the irregular spacing for
205 the Phanerozoic time scale, this criterion is unequally applied compared to the simpler
206 Precambrian age limits set forth in Van der Voo (1990, 1993).

207 Given the many advances in geochronological methods/techniques, we propose a more rigid
208 (and simpler) age constraint on paleomagnetic poles. Our proposal is that the age of the rock
209 (and presumed age of the magnetization) should be known to within ± 15 Ma. Although the
210 blanket ± 15 Ma limit on Phanerozoic rocks allows for a larger percentage error on the younger
211 studies, it is more stringent than the original Q-criterion and allows for reasonable definitions of
212 APWP's.

213 In cases of well-defined remagnetizations, the age criterion should apply to the age of
214 remagnetization rather than that of the rock. For example, a demonstrably synfolding
215 magnetization might be sufficiently well dated if there are independent age constraints on that
216 deformation within the limits of precision set above.

217 Considerable discussion with regard to the age constraints took place amongst the authors
218 with some advocating for more stringent limits. It is important to remember that both the Q and
219 R criteria are not disqualifying. For example, someone investigating rapid true polar wander
220 may want tighter age constraints on relevant paleomagnetic poles and may freely apply their own
221 filter during that analysis.

222 *Reliability Criterion #2- Techniques and Statistical Analyses*

223 The original Van der Voo (1990) criterion #Q2 establishes requirements for measurement
224 precision using Fisher (1953) statistics, whereas criterion #Q3 focused on measurement
225 accuracy. Because imprecise and inaccurate measurements can result from user error or from
226 inadequate sampling of paleosecular variation, we advocate for the following to satisfy R2:

227 a) Attempt at least two methods of stepwise demagnetization (e.g. alternating field and
228 thermal, Meert et al., 1995; or thermal/chemical, Billardello and Kodama, 2011) on at
229 least pilot suite of samples to demonstrate that individual vector components are being
230 separated effectively (figure 4).

231 b) Analyze the directional data using Zijderveld diagrams and principal component
232 analysis (Zijderveld, 1967; Kirschvink, 1980) or great circle intersections to separate

233 overlapping unblocking temperature/coercivity components (Halls, 1976,1978;
234 McFadden and McElhinny, 1988).

235 c) Achieve a VGP scatter that adequately averages paleosecular variation. Methods by
236 which secular variation is assessed may include using a field-based model (McFadden et
237 al., 1991) or a statistical based model (Deenan et al., 2011, 2014), and guided by these
238 approaches we advocate a simple set of statistical tests of the mean result.

239 Fig. 4 here

240 We presuppose that some may view this tripartite list of requirements as being too bulky
241 for one criterion. Whereas sub-criteria (a-b) are now intuitive and routinely performed, the
242 quality of the end result depends on satisfying subcriterion (c). Though not required to satisfy
243 R2, we advocate that authors report several examples of non-ideal behavior when identified.

244 The original sample size, precision parameter and α_{95} requirements for Q2 criteria were
245 somewhat arbitrary. For example, the original α_{95} threshold was chosen empirically as an overlap
246 in error within time constraints set by Q1. Fisher (1953) statistics demonstrate that the number of
247 samples (sites), precision parameter (k) and cone of confidence (α_{95}) are co-dependent. Using the
248 original Q2-criteria of $N = 25$ and $k = 10$ yields an α_{95} of $\sim 9^\circ$ rather than the 16° advocated by
249 Van der Voo (1990). Modern paleomagnetic studies routinely exceed the $N=25$ sample limit
250 proposed by Van der Voo (1990) whereas older studies may not. Therefore, the statistical limits
251 required to satisfy R2 use a different approach (Deenan et al., 2011). By definition, a
252 paleomagnetic pole represents the time-averaged position of the geomagnetic pole that is
253 presumed to be symmetric about the center of the Earth and coaligned with the Earth's spin axis
254 (the Geocentric Axial Dipole, or GAD; Meert, 2009). Debates about the nature of this
255 assumption are beyond the scope of this paper; however, we feel that a quantitative assessment
256 of secular variation should be addressed in any paleomagnetic study (McFadden et al., 1988;
257 Deenan et al., 2011; Tauxe and Kodama, 2009; Lund, 2018).

258 In general, averaging of secular variation is thought to occur over an interval of $\sim 3,000-$
259 $10,000$ years. Sedimentary units should adequately average secular variation over a few meters
260 of sampling (Kodama, 2012), and therefore more easily satisfy this R criterion. On the other
261 hand, quickly cooled igneous rocks provide only a spot reading of the field (McFadden et al.,
262 1988). Secular variation of intrusive rocks is difficult to evaluate as the size of the intrusive
263 body, chemistry of the remanence carriers and temperature of the surrounding country rock

264 affect the timing of remanence acquisition. We prefer the application of a statistical method for
 265 evaluating secular variation as part of the R-criteria.

266 A popular method for evaluating PSV in paleomagnetic studies follows the analysis by
 267 McFadden et al. (1988, 1991) wherein the “Model-G” field (Figure 5; McFadden et al., 1991;
 268 McElhinny and McFadden, 1997) is compared with the observed paleosecular variation (S_T)
 269 using the formula of Cox (1970):

$$270 \quad (2) \quad S_T = \sqrt{\frac{1}{n-1} \sum_{i=1}^n \Delta i^2}; \quad (i=1, \dots, n)$$

271 where n =number of sites (>5) and Δi = angle between the i^{th} VGP and the mean VGP. S_T
 272 represents the sum of geomagnetic secular variation effects S_B and random errors due to
 273 sampling S_W . Nevertheless, in most cases S_T provides a close approximation to S_B .

274 **Fig. 5 here**

275 The quantitative assessment of secular variation noted above applies to studies where
 276 individual sites/samples likely represent a spot reading of the Earth’s magnetic field (basalt
 277 flows, dykes, sills and small intrusions etc. The S_T parameter estimation is still commonly used
 278 to evaluate averaging of secular variation in spite of the fact that there are known mathematical
 279 issues with the model (Tauxe and Kodama, 2009; Deenen et al., 2011; Linder and Gilder, 2012).

280 Deenen et al. (2011, 2014) described in detail the problems in assigning specific k , N and
 281 α_{95} values for the Q2 criterion due to their dependence on one another. For example, 25
 282 *samples* from 3 basalt flows or dykes (cooling units) would meet all the Q2 criterion quite easily
 283 but may be unlikely to provide averaging of secular variation (see also Deenen et al., 2011).
 284 They make the argument that an N -dependent A_{95} (*averaging of virtual geomagnetic poles*)
 285 should be applied to a dataset to satisfy the Q2 criterion. According to their assessment, a
 286 paleomagnetic study where the A_{95} value lies between the following confidence bounds should
 287 provide adequate averaging of secular variation:

$$288 \quad 12 \times N^{-0.40} \leq A_{95_{\text{obs}}} \leq 82 \times N^{-0.63}$$

289 These bounds were established by Deenen et al. (ibid.) to conform to various models of a
 290 time-varying GAD field, but they also serve to demand a minimum expectation of precision for a
 291 valuable paleomagnetic pole. However, the maximum allowable bound on imprecision (A_{95})
 292 may be too lax. For example, using typical values for N among published high-quality
 293 paleomagnetic poles (10-20 sites), the formula yields acceptable limits on Fisher’s K of 1.5–2.6.

294 Such datasets might be marginally acceptable according to some time-varying GAD models, but
295 they are at odds with our experience of reliable poles, which lean more toward what Van der Voo
296 (1990) suggested as having a minimum empirical bound of data clustering ($K \geq 10$). In the other
297 direction, the hallmark signature of a dataset that doesn't adequately average secular variation is
298 *too much* precision on the pole, which would correspond to an anomalously low value of A95
299 and high value of K. When the number of sites is as large as some of the most intensively
300 sampled units ($N \sim 40$ -50 sites), the Deenen et al. (2011) lower bound on A95 corresponds to
301 values of $K \sim 65$ -70. Although smaller sample sets could have ranges of $K > 70$ that conform to
302 statistical GAD models, we suggest that any value of $K > 70$ warrants some suspicion of
303 inadequate averaging of secular variation.

304 In addition to the requisite bounds on K between 10 and 70, we also suggest a minimum
305 number of independent spot readings of the ancient magnetic field. Van der Voo (1990) cited
306 personal experience in assigning $n \geq 25$ samples for a pole's reliability, and we broadly accept
307 that order of value. Multiple samples should be collected from each site in order to average
308 within-site or between sample errors (McElhinny and McFadden, 2000). Opdyke and Channell
309 (1996) suggest that three or more samples are required for unambiguous determination of
310 polarity at each site. Therefore, we suggest that the 'test' for PSV should be applied to a study
311 with $N \geq 25$ (samples), $10 \leq K \leq 70$ and $B \geq 8$ sites (a site represents a spot reading of the
312 magnetic field; *minimum of 3 samples per site*).

313 In summary, meeting the R2 criterion requires both adequate demagnetization and
314 sampling to achieve the goal of averaging secular variation. As a final note, the authors of this
315 proposal, as both users and developers of paleomagnetic databases, appeal to the community at-
316 large to consider including, at a minimum, a specific set of information in each publication that
317 includes new results, listed by example in Table 2. Inclusion of this information in each
318 publication facilitates entry into global databases and evaluation of R-criteria. Authors may feel
319 free to add more entries into their data tables, but the data shown in example Table 2 are
320 essential for database compilations and other calculations.

321 **Table 2 Here**

322 *Reliability Criterion #3- Characterization of Magnetic mineralogy/rock magnetism*

323 Modern studies should include an investigation into the magnetic carriers via rock
324 magnetic tests and/or petrographic examination. Characterization of the magnetic carriers aids

325 in determining the primary/secondary nature of a particular remanence direction (Halls and
326 Zhang, 1995; Halls et al., 2001; Jackson and Swanson-Hysell, 2012; Auborg et al., 2012;
327 Zechmeister et al., 2012; Kodama and Dekkers, 2004).

328 A general description of magnetic carriers includes an evaluation of any of the following:
329 unblocking/coercivity spectra, isothermal remanent acquisition (IRM) tests, temperature-
330 susceptibility analyses, 3-axis IRM (Lowrie, 1990), low-temperature treatment of IRM (Nagata
331 et al., 1964; Ozdemir et al., 1993; Dekkers et al., 1989). Hysteresis properties are useful in
332 evaluating the domain size of remanence carriers. Hysteresis studies may include Day plots
333 (Day et al., 1977; Roberts et al., 2018) and first order reversal curves (FORC diagrams; Pike et
334 al., 1999; Roberts et al., 2000). In addition, magnetic fabric studies have proven useful in
335 evaluating remanence carriers and/or deformation. These include anisotropy of magnetic
336 susceptibility (AMS; Graham, 1954, 1957), anisotropy of isothermal remanence (AIR; McCabe
337 et al., 1985) and anisotropy of anhysteretic remanence (AAR).

338 Microscopic investigations of magnetic carriers using polished thin sections under
339 reflected light; scanning electron or transmission electron microscopes help identify possible
340 magnetic carriers and their relationship to the original petrology (Poldervaart and Gilkey, 1954;
341 Pichamuthu, 1959; Halls and Zhang, 1995; Halls et al., 2007; Sun and Jackson, 1994). In the
342 case of fine-grained sediments, mineral separation techniques may be applied to identify the size
343 and composition (Opdyke and Channell, 1996).

344 The identification of magnetic carriers is particularly important in clastic sedimentary
345 rocks (redbeds) where inclination shallowing during detrital remanence acquisition (DRM) can
346 adversely affect tectonic interpretations (King, 1955; Gilder et al., 2001; Tan and Kodama, 2003;
347 Tauxe and Kent, 2004; Li and Kodama, 2016). Chemical remanent magnetization (CRM) can
348 post-date, and overprint, depositional remanence (DRM) through a significant time interval,
349 complicating paleomagnetic interpretations (Kodama and Sun, 1992; Kodama and Dekkers,
350 2004; Jiang et al., 2015).

351 R3 criterion will be met if there is a reasonable attempt to identify and comment on the
352 significance of the magnetic carriers in the study, either through petrographic or rock-magnetic
353 investigation.

354

355 *Reliability Criterion #4- Field Tests that constrain the age of magnetization*

356 a. Baked Contact/Inverse Baked Contact Test

357 A study can receive the R4 criterion provided that the baked contact test exhibits most of
358 the features outlined in Figure 2a ($R4_{C+}$). The baked contact test often departs from the ideal
359 models described above. A positive baked contact ($R4_{C+}$) test is also confirmed when the dyke
360 and baked zone exhibit stable and similar paleomagnetic directions and the unbaked region
361 yields a stable and different direction even if there is no hybrid zone. It is not uncommon for the
362 unbaked host rock to exhibit unstable behavior regardless of whether or not regional
363 remagnetization has occurred. In the case where the intrusive body and baked host show the
364 same direction and the country rock exhibits unstable behavior, the baked contact test should be
365 noted as $R4_{Co}$. Salminen et al. (2009) note a special case where heating associated with meteorite
366 impact may provide evidence of a primary magnetization in the melt zone and adjacent regions.
367 Inverse baked contact tests, as long as they satisfy the characteristics noted above, may also
368 provide useful age constraints on magnetization directions in the host rocks and would also
369 qualify for R4.

370 b. Fold/Tilt/Slump Tests

371 The strongest fold tests are those that (a) pass rigorous statistical analyses and (b)
372 have an age of folding that is ‘close’ to the age of the rocks in question. The fold test should be
373 applied in a stepwise manner and we require that the magnetization direction have optimal
374 grouping within error of 90-110% unfolding in order to meet R4 standards. Although the fold
375 test provides more clarity on the age of magnetization when the age of folding is close to the age
376 of the rocks (Van der Voo, 1969), the R4 positive fold test would be satisfied regardless of the
377 age of folding. Regarding the question of whether a pole is reliable for a given purpose, we
378 prefer to leave this decision in the hands of each individual in the context of their particular
379 analysis. In similar fashion, we recognize several statistical variations of the fold test, catered to
380 a variety of fold geometries and sampling strategies; any statistically robust test can be used to
381 satisfy R4.

382 The intention of R4 is to identify evidence in favor of the possibility of a primary
383 magnetization in the rock. Thus, syn-folding magnetizations do not meet the R4 criterion unless
384 they are demonstrably syn-sedimentary slump folds (Smith et al., 1983; Schmidt et al., 1991).
385 We do not devalue the significance of a syn-folding (re)magnetization, but since the result
386 neither ‘passes’ nor ‘fails’ the fold test, we prefer a simplistic approach in the grading scheme.

387 Individual researchers may pass judgement on the validity of a syn-folding magnetization as
388 needed

389 c. Conglomerate Test

390 A paleomagnetic pole will receive the R4 criterion for the conglomerate test if it (a)
391 fulfills

392 the statistical requirements set forth in Watson (1956b), Shipunov et al. (1998) or Heslop and
393 Roberts (2018a) and (b) N is sufficiently large to test the null hypothesis H_0 of Watson (1956b)
394 which assumes a uniform (“random”) distribution of vectors or the Bayesian assumptions set
395 forth in Heslop and Roberts (2018a). Heslop and Roberts (2018a) tested sample sizes ranging
396 from $n=5$ to $n=35$. While not specifically assigning an optimal N-value, they note that a strong
397 level of support for the conglomerate test is difficult for sample sizes where $n < 19$. We attempt
398 to balance statistical vagaries with practicality in field sampling and argue that $n \geq 10$ in order to
399 be a useful conglomerate test. Providing that the statistical analysis indicates a positive
400 conglomerate test, the age of the conglomerate should, in principle, be reasonably close to the
401 age of the rocks being studied; however, as with the fold test described above, we do not place a
402 specific restriction on the age of a conglomerate (other than devaluing any misuse of the
403 reliability scale by applying the test to trivially young conglomerates or breccias). The ideal case
404 requires that the conglomerate clasts are taken from an intraformational conglomerate wherein
405 the clasts are derived from the underlying units and exhibit the same magnetic characteristics as
406 the parent materials (see for example Buchan and Hodych, 1989; Levashova et al., 2009; Meert
407 et al., 1994; Meert et al., 2009).

408 Fig 6 Here

409 d. Unconformity Test

410 The unconformity test (Kirschvink, 1978) was proposed for the special case in which
411 a stratigraphically ordered polarity sequence is truncated by an unconformity. Figure 6a
412 illustrates a positive unconformity test wherein the polarity sequence below the unconformity is
413 discontinuous across the unconformable surface. In this case, the magnetization in the lower
414 sequence is older than the unconformity. Figure 6b illustrates a negative unconformity test
415 because the polarity zonation is continuous across the unconformity.

416 *Reliability Criterion #5- Structural control*

417 We accept the original rationale employed by Van der Voo (1990) and argue that
418 paleomagnetic poles derived from allochthonous or parautochthonous terranes, non-stratified
419 rocks and regions that have undergone internal vertical axis rotations will not meet the R5
420 criterion. Results from intrusive rocks younger than the last deformational event may meet this
421 criterion. We also note that we apply the R5 criteria in a stricter fashion than Van der Voo
422 (1993). As an example, in the VdV (1993) compilation of paleomagnetic poles from Laurentia,
423 results from the Colorado Plateau and northern limb of the Pennsylvania salient were ‘corrected’
424 for vertical-axis clockwise rotations of 5 and 23 degrees respectively and each pole from those
425 regions received a Q=5 checkmark. The tacit assumption was that the amount of rotation for
426 each region was well-known; however, the magnitude of the CP rotation and ‘corrections’ turn
427 out to be far more complicated (see McCall and Kodama, 2014). Therefore, a pole will meet R5
428 if there is a presumption that the region was a rigid part of the craton since the time the
429 magnetization was acquired.

430 The definition of autochthonous in the Precambrian is complicated by the fact that
431 Phanerozoic ‘continents’ are themselves amalgams of smaller nuclei with a complex assembly
432 history (Kilian et al., 2017; Hoffman, 1998; Meert and Pandit, 2015; Bogdanova et al., 2008;
433 Gladkochub et al., 2006; Cawood and Korsch, 2008; Boger, 2011; de Waele et al., 2008;
434 Tassinari and Macambira, 1999). Most modern paleomagnetic studies recognize this obstacle
435 and use poles that represent only the region to which they are rigidly attached (our stipulation).
436 The situation is sometimes further complicated by “post-assembly” rotation or rifting. For
437 example, Cawood and Korsch (2008) argue that three key elements of Australia (Northern
438 Australia, Western Australia and the Mawson continent) were assembled during the
439 Mesoproterozoic. Li and Evans (2011) provide a convincing argument for a large 40° degree
440 intraplate rotation between a coupled western Australia/southern Australia and northern Australia
441 during the late Neoproterozoic. In this example, Mesoproterozoic poles calculated from the
442 blocks that have undergone rotation can receive the R5 criterion when applied to their respective
443 regions. Authors assigning R5 to poles referred by their studies should specify their definition of
444 each “craton” along with its present and paleogeographic bounds.

445 Clastic sedimentary rocks often carry a detrital remanent magnetization (DRM) that can
446 experience inclination flattening during deposition and compaction (King, 1955; Figure 7). The
447 relationship between flattening factor (f) and inclination is given as:

$$f \tan(I_f) = \tan(I_0)$$

448 Where:

449 f = flattening factor ($0 \leq f \leq 1$)

450 I_f = expected GAD inclination for latitude of deposition

451 I_0 = observed inclination

452 **Fig 7 here**

453 There are two main strategies employed for detecting and correcting inclination
 454 shallowed directions (Jackson et al., 1991; Kodama and Sun, 1992; Kodama and Dekkers, 2003;
 455 Kodama, 1997; Li and Kodama, 2016; Tauxe and Kent, 2004; Tauxe et al., 2008). The first
 456 technique is based on detailed measurements of anisotropy (Li and Kodama, 2016) that are labor
 457 intensive but may provide a more direct measure of inclination shallowing than the ‘easier’
 458 statistical analysis (Tauxe and Kent, 2004). As noted by Li and Kodama (2016), both
 459 techniques have underlying assumptions and limitations that require a cautious approach in
 460 drawing paleogeographic conclusions from inclination-shallowed rocks. We therefore argue that
 461 any poles based on flattening corrections will not meet the R5 criterion unless the inclinations
 462 are corroborated by paleomagnetic data from either coeval volcanic rocks that have a similar R-
 463 value or other sedimentary rocks that do not require flattening corrections.

464 *Reliability Criterion #6- Presence of Reversals- Statistically valid reversal test (McFadden and*
 465 *McElhinny, 1990; Heslop and Roberts, 2018b)*

466 The power of the reversal test in paleomagnetism is based on the assumption that a
 467 positive result indicates sufficient passage of time required to average secular variation.
 468 Furthermore, antipodal directions suggest that there were no systematic overprints on the
 469 primary magnetization.

470 Unfortunately, the reversal test has sometimes led to the false conclusion that the rocks
 471 record a primary magnetization. This is neither the intended purpose of a reversal test nor
 472 always an accurate assumption. Dual-polarity remagnetization is possible (see Johnson et al.,
 473 1984; Johnson and Van der Voo, 1989) and single polarity results can be demonstrably primary.
 474 Furthermore, data collected from ‘spot’ readings of the geomagnetic field (smaller dykes or flow
 475 units) may exhibit dual polarity magnetization without adequately averaging secular variation. A
 476 positive reversal test is therefore supportive, but not conclusive, of a PSV-averaged primary
 477 magnetization in the sampled sequence (see R2).

478 In the original Q-criteria compilation (Van der Voo, 1990) there was no robust statistical
 479 test required for meeting this criterion other than the presence of both polarities with overlapping
 480 α_{95} confidence limits. In part, this was because the statistical test proposed by McFadden and
 481 McElhinny (1990; hereafter M&M) had not been sufficiently applied to the extant database. The
 482 reversal test of M&M (1990) grades significance by comparing the means of the normal and
 483 reverse directions assuming that they are drawn from the same population. The assumption of a
 484 common population depends on the number of observations of each polarity and the precision
 485 parameter kappa (k). If there is no common precision parameter, or isolated observations from
 486 one of the polarity groups, then the test is not necessarily invalid, but requires additional analysis
 487 (see M&M, 1990). We note that the sample size and common kappa assumption for the M&M
 488 (1990) test have a key effect on the test results. If the common population assumption is
 489 statistically valid (i.e. $\gamma_o < \gamma_c$) then the reversal test is evaluated according to the critical angle
 490 (γ_c). A positive reversal test is graded “R_A” when $\gamma_c < 5^\circ$; “R_B” when $\gamma_c < 10^\circ$; “R_C” when $\gamma_c <$
 491 20° and “Indeterminate”, or “R₀”, when $\gamma_c \geq 20^\circ$. A negative reversal test occurs when $\gamma_o > \gamma_c$.
 492 If the test is based on isolated observations from one of the polarity groupings, we propose
 493 following the suggestion of M&M (1990) of assigning grades R_{AI}, R_{BI} and R_{CI} for the reversals
 494 test.

495 A second reversal test was recently proposed by Heslop and Roberts (2018b; hereafter
 496 H&R) that is more nuanced in grading the reversals test. The H&R (2018b) test is particularly
 497 useful when the M&M (1990) result is indeterminate (R₀). H&R (2018b) show that ~40% of
 498 “R₀” reversal tests under the M&M (1990) scheme yield positive support for a common mean,
 499 ~59% are ambiguous and <1% yield positive support for a different mean. To meet the R5
 500 criterion, a positive reversal test must rise above the M&M (1990) “indeterminate” or the H&R
 501 (2018b) “ambiguous” label. As a cautionary note, we note that the SIAPD program (Torsvik et
 502 al., 1999) reversal test assumes the user has evaluated the common kappa and sample size
 503 requirements for the M&M (1990) test. These assumptions must be met in order to avoid
 504 incorrect calculation of reversal test results (Nagaraju et al., 2018; Kumar et al., 2017).

505 *Reliability Criterion #7: No Resemblance to younger poles by more than a period unless there is*
 506 *field evidence for an older magnetization*

507 One of the more contentious discussions in Precambrian paleomagnetism is whether or
 508 not Q7 “resemblance to a younger paleopole by more than a period” should be considered in the

509 revised “R” reliability criteria. Veikkolainen et al. (2014), during their compilation of the global
 510 Precambrian database, argued that Q7 should be disregarded because it could lead to erroneous
 511 conclusions about remagnetization. They argued that the presence of several self-closing APWP
 512 loops in the Precambrian were merely coincidental and not indicative of remagnetization.

513 Using an argument nearly antithetical to that of Veikkolainen et al. (2014), Bazhenov et
 514 al. (2016) took a pessimistic approach towards the Precambrian database from Baltica.
 515 Bazhenov et al. (2016) calculated the statistical probabilities for a 95% confidence deviation
 516 from the true mean (see McFadden, 1980; $p=.05$, N =number of observations; R =length of the
 517 resultant mean vector):

$$518 \cos\psi_{(1-p)} = 1 - (N - R) \left[\left(\frac{1}{p} \right)^{\frac{1}{N-1}} - 1 \right]$$

519 The approximation for the 95% deviation angle is given by:

$$\psi_{95} = \frac{140}{\sqrt{k}}$$

520 ψ_{95} was then used to create a double-width ‘alarm band’ around the Phanerozoic APWP for
 521 Baltica (Figure 8; using an average k of 100 from published studies). They observed that ~50%
 522 of Precambrian poles; (a) fell within the ‘alarm band’; (b) were not randomly distributed and; (c)
 523 formed distinct clusters which they concluded should be viewed with a suspicion of
 524 remagnetization.

525 **Fig. 8 here**

526 Pivarunas et al. (2018) approached the issue by generating hundreds of synthetic APWP’s
 527 in an effort to evaluate the statistical probability of self-intersection (Figure 9). Pivarunas et al.
 528 (2018) show that the likelihood of APWP self-intersection is $69.1 \pm 9\%$ in 500 million years and
 529 $97.1 \pm 2\%$ in 1000 million years with a ‘plate-reorganization’ every 70 Ma. In other words,
 530 resemblance to younger paleopoles noted by Bazhenov et al. (2016) and Veikkolainen et al.
 531 (2014) are the expected outcome of continental motion over geological time.

532 **Fig 9 Here**

533 There are disagreements regarding inclusion of this criterion in our revision amongst the
 534 authors of this paper. Some side with the more conservative approach of Bazhenov et al. (2016)
 535 whilst others favor the abolition of this criterion. Arguments for abandoning this criterion were
 536 summarized by one of us (D.A.D. Evans) as follows:

537 (1) The definition of ‘resemblance’ is not clearly defined in the Van der Voo (1990)
538 criterion.

539 (2) What level of reliability is required of the younger pole?

540 (3) How far do we draw our geographic boundaries when attempting to assess
541 ‘resemblance’ for a particular region? This is particularly relevant when assessing
542 Precambrian poles.

543 These are legitimate and confounding issues. In some cases, the solution is simple. Any
544 paleomagnetic pole with field tests that constrain the age of magnetization to be older than the
545 younger pole(s) it resembles will meet this criterion.

546 Dealing with the other concerns raised above is more problematic. Van der Voo (1990)
547 stated that any pole that fell on a younger part of the APWP should be viewed with suspicion.
548 The implication is that a cratonic block must have a well-defined path for comparison, but few
549 cratons have well-defined Precambrian APWP’s and in some cases the age of cratonic coherence
550 is not well-established (Meert and Pandit, 2015). Furthermore, no specific guidance was offered
551 by Van der Voo (1990, 1993) regarding points (1) and (2).

552 We offer the following instructions for evaluating R7.

553 (1) Comparison to younger poles: Heslop and Roberts (2019) discuss the difficulty in
554 assessing what constitutes ‘resemblance to a younger pole’. They argue that a binary
555 “yes’ or ‘no’ decision is difficult and propose a series of information metrics that can aid
556 in this decision. In spite of known limitations outlined in that study, we propose that
557 paleomagnetic poles with overlapping A95 envelopes with younger poles (of $R \geq 3$) will
558 not meet this criterion. Individual investigations may wish to apply the metrics described
559 in Heslop and Roberts (2019), but we choose to apply the more conservative approach of
560 overlapping A95 confidence intervals for our criterion. This approach may be justified
561 at least qualitatively by the recognition that A95 values indicate statistical precision but
562 do not always represent all possible sources of error in paleomagnetic data (e.g., Rowley,
563 2019).

564 (2) Geographic boundaries: The comparison to younger poles should only be made with
565 poles from stable regions within the connected craton(s). Poles from orogenic belts
566 should not be used for this comparison as they do not provide a unique pole position.

567 This requires knowledge of the assembly history the continent/craton being evaluated and
 568 is best considered by the authors at the time of publication.

569 There is no perfect solution to the issue of remagnetization, but we feel that Van der Voo's
 570 (1993) cautionary statement "guilty (i.e. remagnetized) until proven innocent" is still valid.

571 4. Conclusions & Recommendations

572 The Van der Voo (1990) Q-criteria served the paleomagnetic community for nearly 30 years
 573 and resulted in more careful and detailed paleomagnetic studies. Modern paleomagnetic
 574 methods, automation, advanced statistical tools along with better precision in geochronological
 575 methods necessitated a re-evaluation of the reliability criteria and our proposal for the new "R"
 576 system. Similar to its predecessor, the R-factor is based on seven criteria used to assess the
 577 reliability of a paleomagnetic pole. These seven criteria are presented in Tabular form (Table 3).
 578 We emphasize that the R-criteria merely form a checklist that provides a numerical reliability
 579 score. The R-score value does not imply rejection or endorsement of any individual
 580 paleomagnetic study. Decisions as to how to apply the R-score to a particular study is up to the
 581 individual researcher or research group.

582 [Table 3 here](#)

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 588

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897 Figure Legends

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- 900 **Figure 1.** Conceptual diagram of (a) positive fold test correction, where rotating inclination and declination data
 901 back to pre-folded values results in agreement between data within the folded bed and; (b) post-folding
 902 magnetization, where rotating inclination and declination back to pre-folded values results in scattering of data from
 903 within the folded unit.
- 904 **Figure 2.** (a) expected unblocking temperature spectra (T_{ub}) along a baked profile showing completely baked host,
 905 hybrid host and unbaked host; (b) Stereonet showing an idealized positive baked contact test (C+); (c) Stereonet of
 906 vector components in an inconclusive baked contact test (Co). Although the baked host matches the dyke direction,
 907 results away from the bake zone do not exhibit stable hybrid or stable host directions. (d) Negative baked contact
 908 test where all directions are similar suggesting a more widespread remagnetization.
- 909 **Figure 3.** (a) positive intraformational conglomerate test (from Meert et al., 2009). Bounding layers show a stable
 910 magnetization above and below the conglomeratic layer. Clasts from the intervening layer exhibit the same
 911 demagnetization behavior as their parent materials with statistically random directions (b) negative intraformational
 912 conglomerate test where layers above and below the conglomerate show identical directions and clasts from the
 913 conglomerate are clustered in the same location.
- 914 **Figure 4.** (a) Stereoplot of alternating field (AF-top) versus thermal demagnetization (bottom) directional changes
 915 from a Paleoproterozoic dyke in India. (b) Plot of thermal (to 560 C) followed by AF-demagnetization of a
 916 Paleoproterozoic dyke in India. In both (a) and (b) thermal demagnetization is ineffective in isolating the
 917 characteristic NW-shallow up magnetization (Pivarunas et al., in prep). (c) Alternating field demagnetization of
 918 Mbozi Complex and (d) Thermal demagnetization of Mbozi Complex. In this case, AF-demagnetization was unable
 919 to resolve the characteristic direction (Meert et al., 1995).
- 920 **Figure 5.** Model G (McElhinny and McFadden, 1997) based expected VGP scatter at different latitudes using
 921 equation (1). For example, a mean VGP scatter of 20 degrees would be expected at a latitude of 60 degrees.
- 922 **Figure 6.** (a) Positive unconformity test (after Kirschvink, 1978). Reversal pattern is truncated across the
 923 unconformity surface; (b) Negative unconformity test. Reversal pattern is continuous across the unconformity
 924 surface.
- 925 **Figure 7.** Effect of inclination shallowing versus geocentric axial dipole inclination for a range of flattening factors
 926 ($f=0.3$ orange; $f=0.6$ grey; $f=0.8$ yellow; GAD field blue).

927 **Figure 8.** From Bazhenov et al. (2016). Alarm band (light green) surrounding the Phanerozoic apparent polar
928 wander path for Baltica along with Precambrian poles from Fennoscandia, Ukraine and the Urals.

929 **Figure 9.** From Pivarunas et al. (2018) (a) example of self-intersecting APWP and (b) a 500 Myr long non-
930 intersecting APWP. Both (a) and (b) use randomly generated velocities between 2-10 cm/yr with a plate re-
931 organization interval of 70 Ma.

Table 1. Van der Voo (1990) Criteria Summary

Q	Brief Description	Limits
1	Well-determined rock age and a presumption that magnetization is the same age	Within ½ period or ± 4% (whichever is larger) in the Phanerozoic. +/-4% or 40 Ma (whichever is smaller) in the Precambrian
2	Sufficient number of samples and statistical limits	$k(K) \geq 10$, $\alpha_{95}(A_{95}) \leq 16^\circ$, $N \geq 25$ samples
3	Adequate demagnetization that demonstrably includes vector subtraction.	Zijderveld (1967), PCA (Kirschvink, 1980) or great circle analyses (Halls, 1976, 1978)
4	Field Tests that constrain age of magnetization	Positive fold, baked contact or conglomerate tests that are statistically valid.
5	Structural control, and tectonic coherence with craton or block involved	Results from thrust sheets or intrusives older than the last tectonic phase not valid.
6	The presence of reversals	Presence of dual-polarity magnetization. No test required
7	No resemblance to paleopoles of younger age (by more than a Period)	No suspicion of remagnetization.

Table2

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Table 2. Sample Data Table for Paleomagnetic Results

Site	Slat	Slong	N/n	Dec	Inc	α_{95}	k	VGP Lat	VGP Long
I915	25.1232° N	87.3456° E	7/7	45°	+55°	8°	56	47.7° N	151.3° E
I916	25.1342° N	86.9984° E	8/6	53°	+47°	7°	112	42.9° N	160.9° E
I917*	24.9765° N	87.0034° E	7/3	123°	+14°	27°	11	26.0° S	154.8° E
I918	24.9965° N	87.1254° E	9/8	48°	+52°	5°	78	47.6° N	155.3° E
Mean Result	Specify mean of VGP's or mean D,I and reference locality		Specify whether mean is based on unit samples (n) or unit sites (B). Specify any data NOT used in calculating mean.	49°	+51.4°	7.3°	288	46.1° N A95=6.6°	156.0° E K=348

Slat=Site latitude; Slong=Site Longitude; N=samples; n=samples used; Dec=Mean Declination; Inc=Mean Inclination; α_{95} =cone of 95% confidence about the mean result; k=Fisher precision parameter; VGP Lat=virtual geomagnetic pole latitude; VGP long=virtual geomagnetic pole longitude; A95= cone of confidence about the mean paleomagnetic pole; K= Fisher's precision parameter in pole space; *Site not used to calculate mean result (Please make sure rejected sites are properly annotated).

Table 3. Brief Description of the R-Score

R	Brief Description	Limits
1	Well-determined rock age and a presumption that magnetization is the same age	Radiometric age constrained to within +/-15 Ma
2	Techniques and Statistical analysis	Stepwise demagnetization effectiveness confirmed by multiple demagnetization methods. Test for averaging of PSV. $N \geq 25$, $10 \leq K \leq 70$, $B \geq 8$ sites (minimum 3 samples/site)
3	Evaluation of remanence carriers	Rock magnetic and/or microscopic examination and identification of magnetic carriers.
4	Field Tests that constrain age of magnetization	Fold/tilt test. Baked contact tests; conglomerate test or other field tests that constrain age of magnetization.
5	Structural control, and tectonic coherence with craton or block involved. Inclination shallowing assessed in clastic sedimentary rocks	Data from thrust sheets or intrusive rocks must be younger than the last tectonic deformation in the area. Detrital sedimentary rocks that do not require inclination corrections will meet this criteria.
6	The presence of magnetic reversals	Statistically significant antipodal normal and reverse directions R_a , R_b or R_c rated (McFadden and McElhinny, 1990) or show support for a common mean (Heslop and Roberts, 2018b).
7	No resemblance to younger poles by more than a period based on overlapping A95	Field tests that constrain the magnetization to be older than the younger pole(s) it resembles.

Figure 1

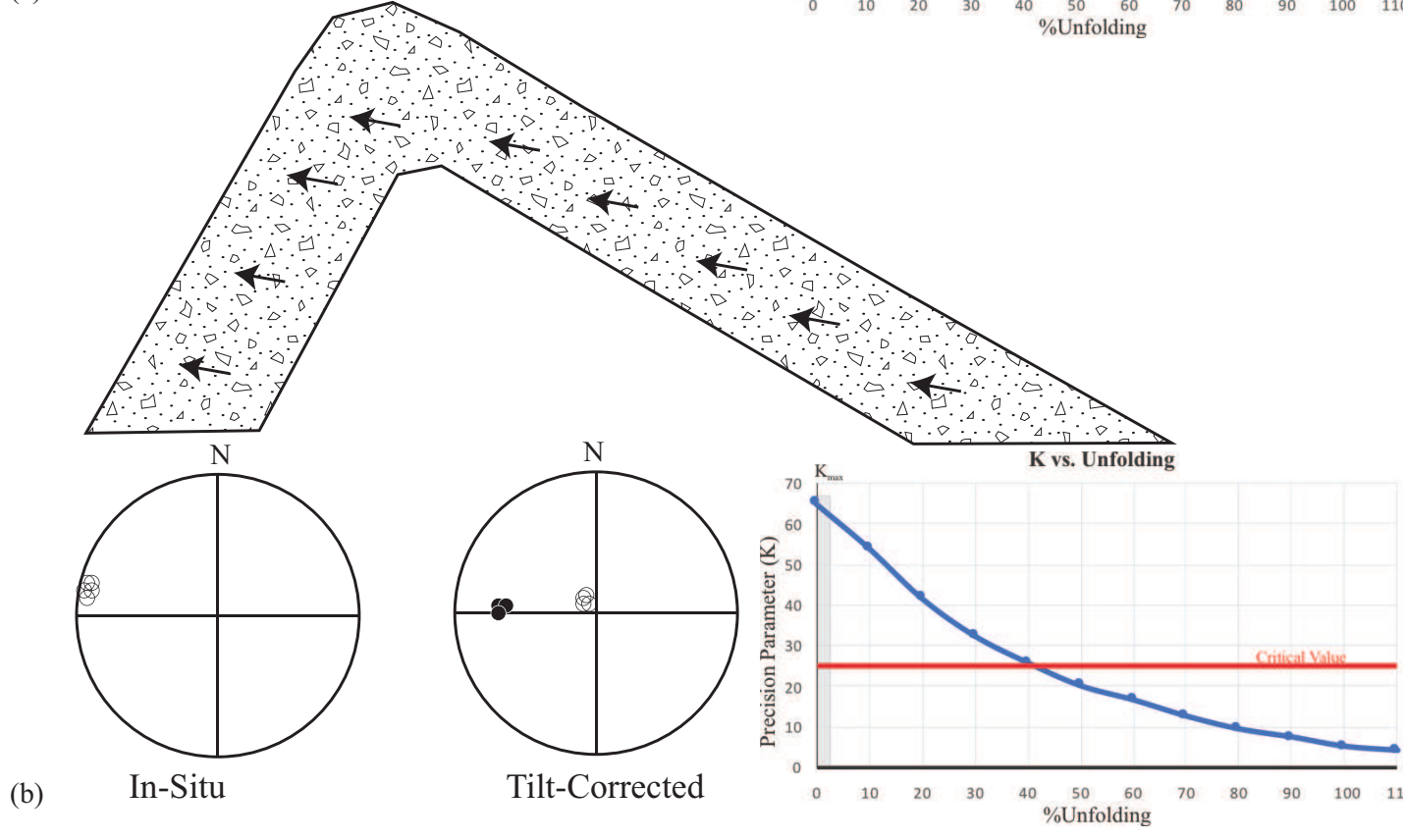
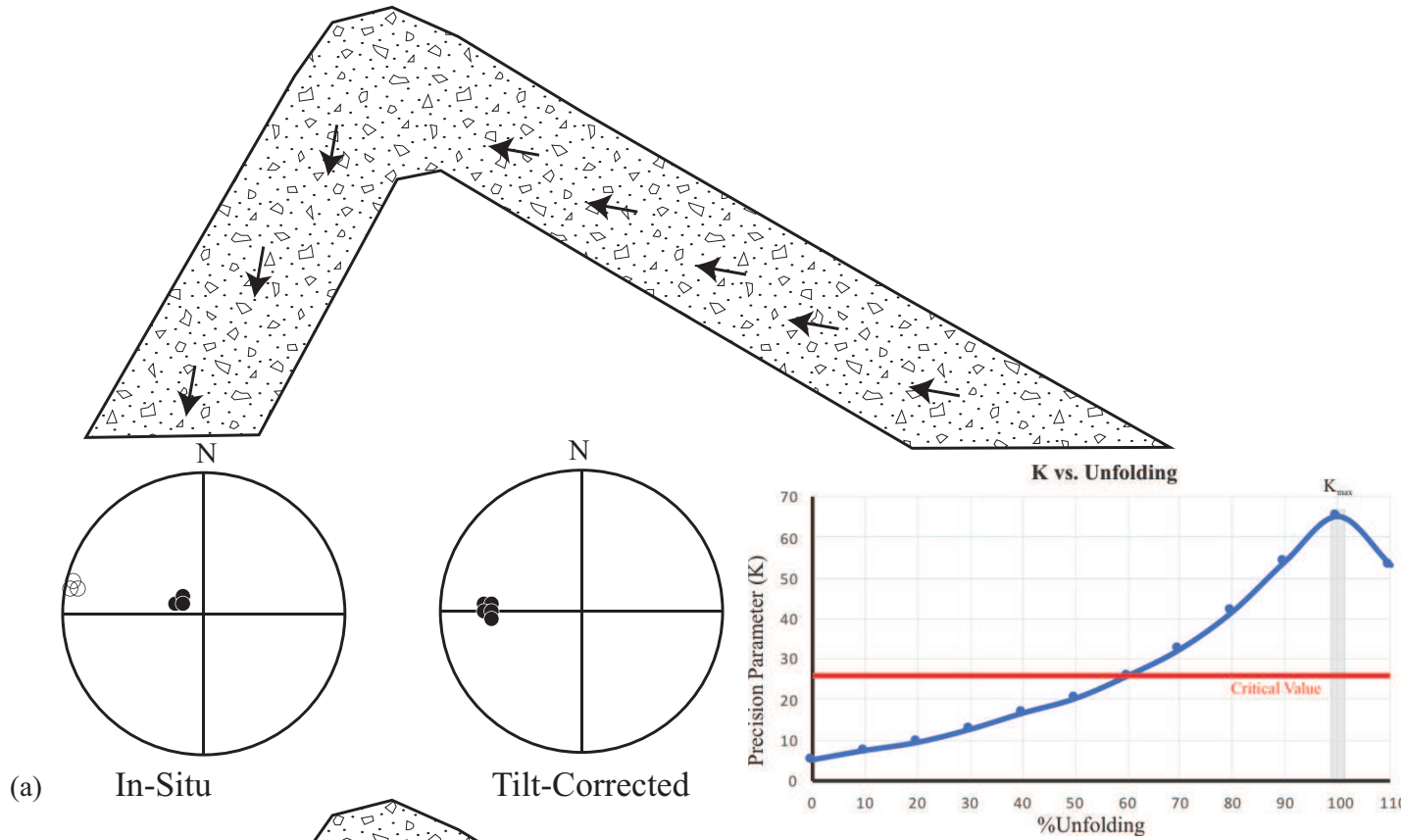
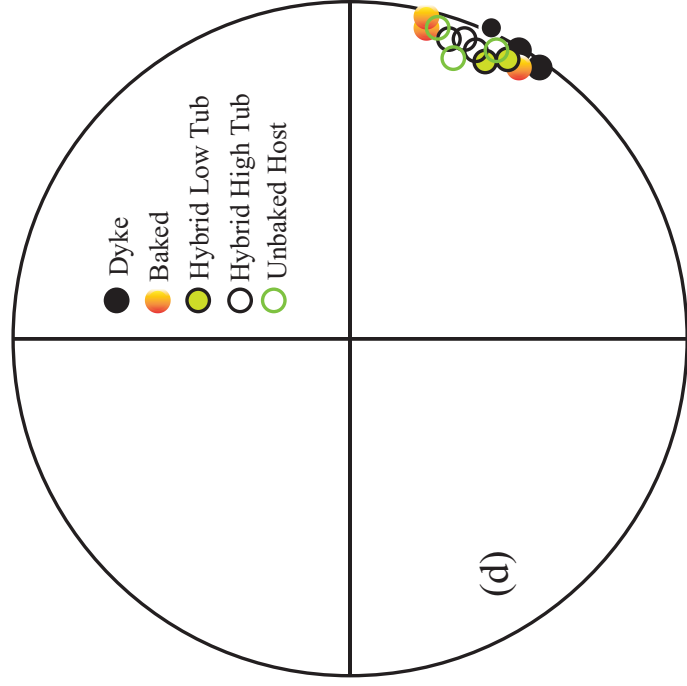
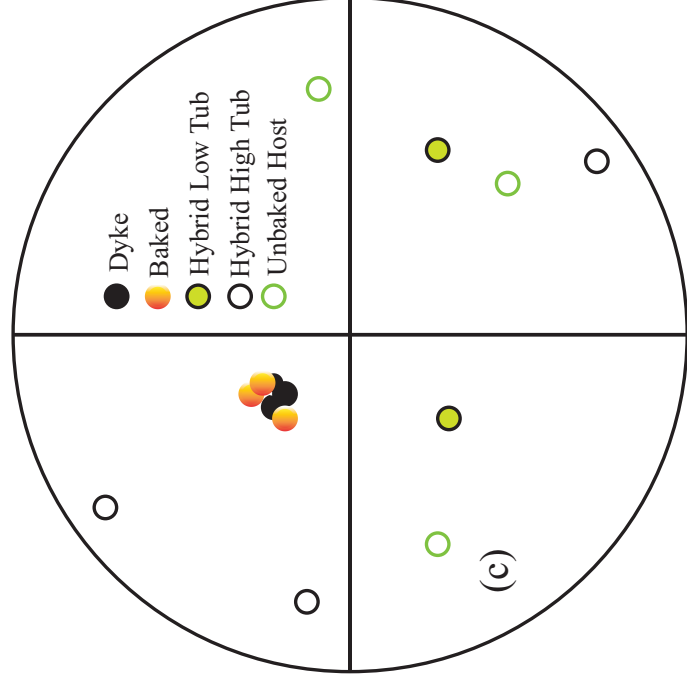
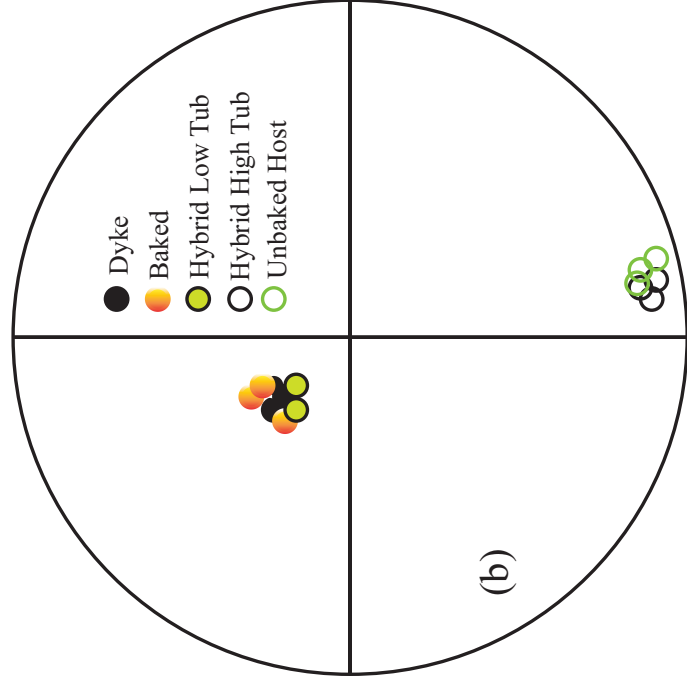
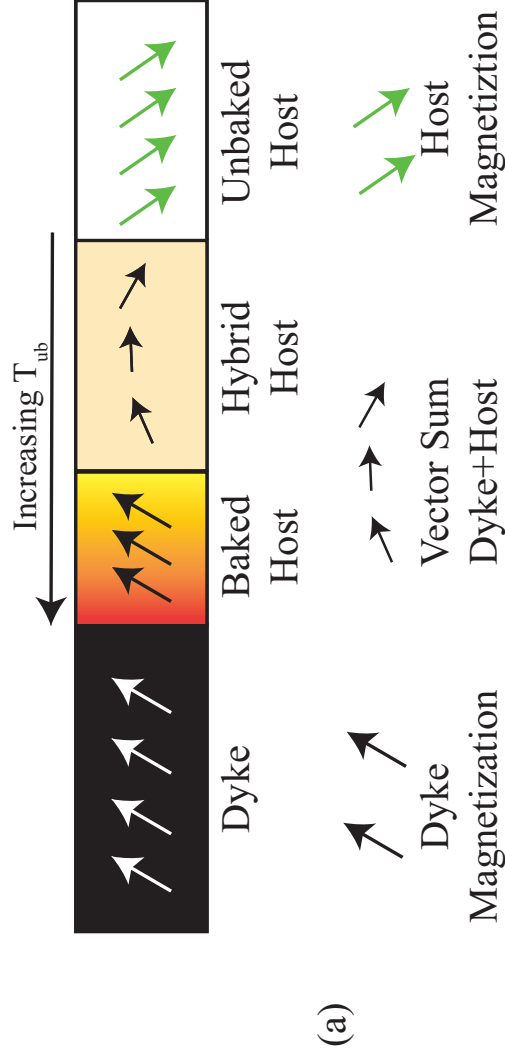


Figure2



Positive Baked Contact Test

Inconclusive Baked Contact Test

Negative Baked Contact Test

Figure3

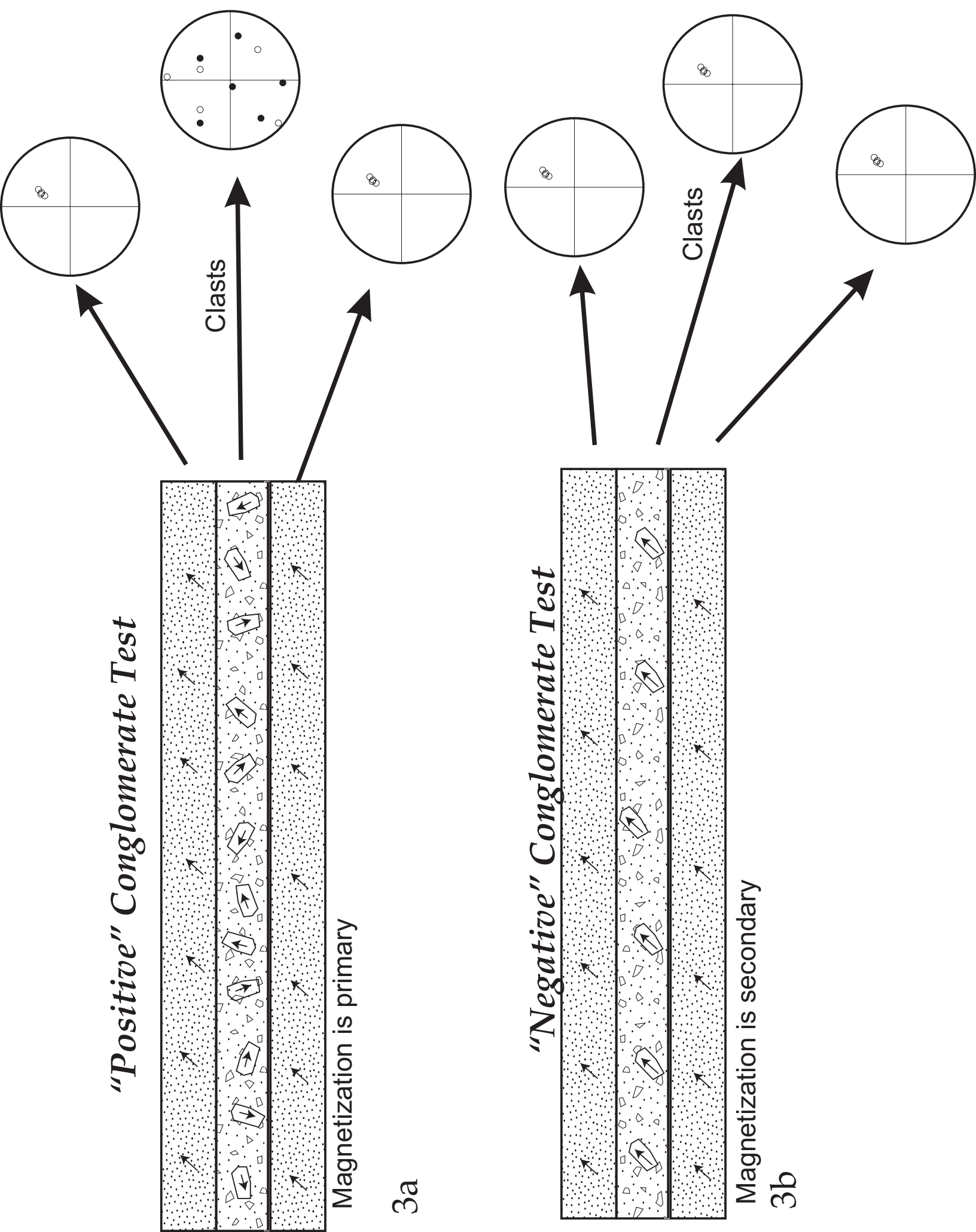


Figure 4

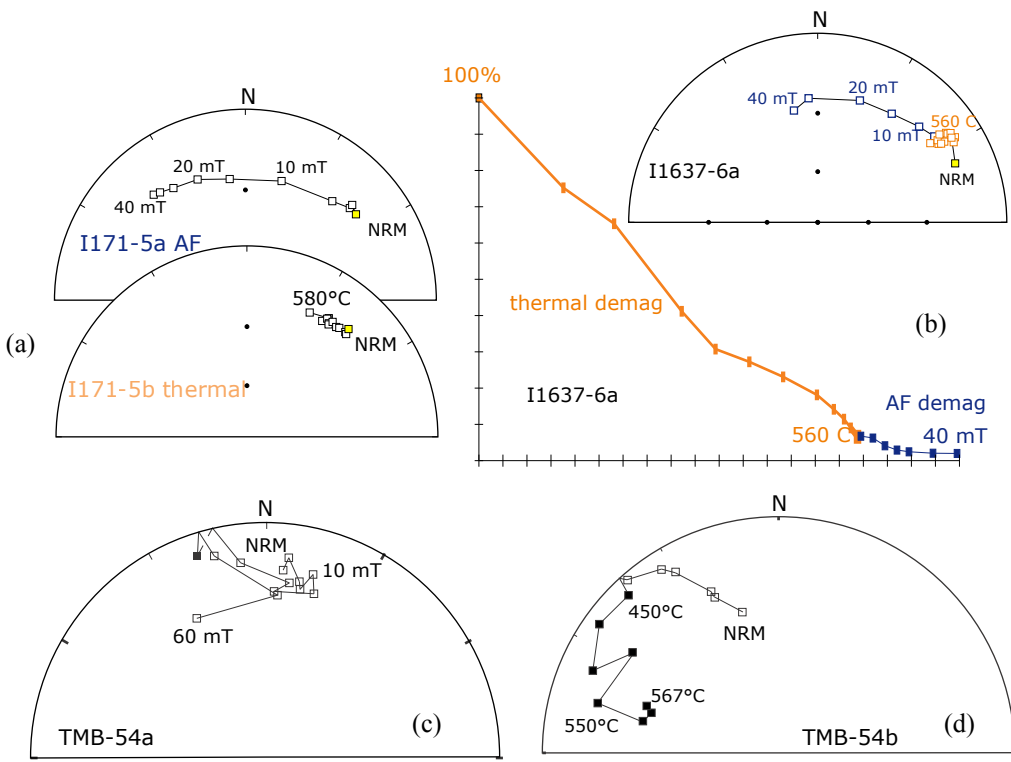


Figure 530

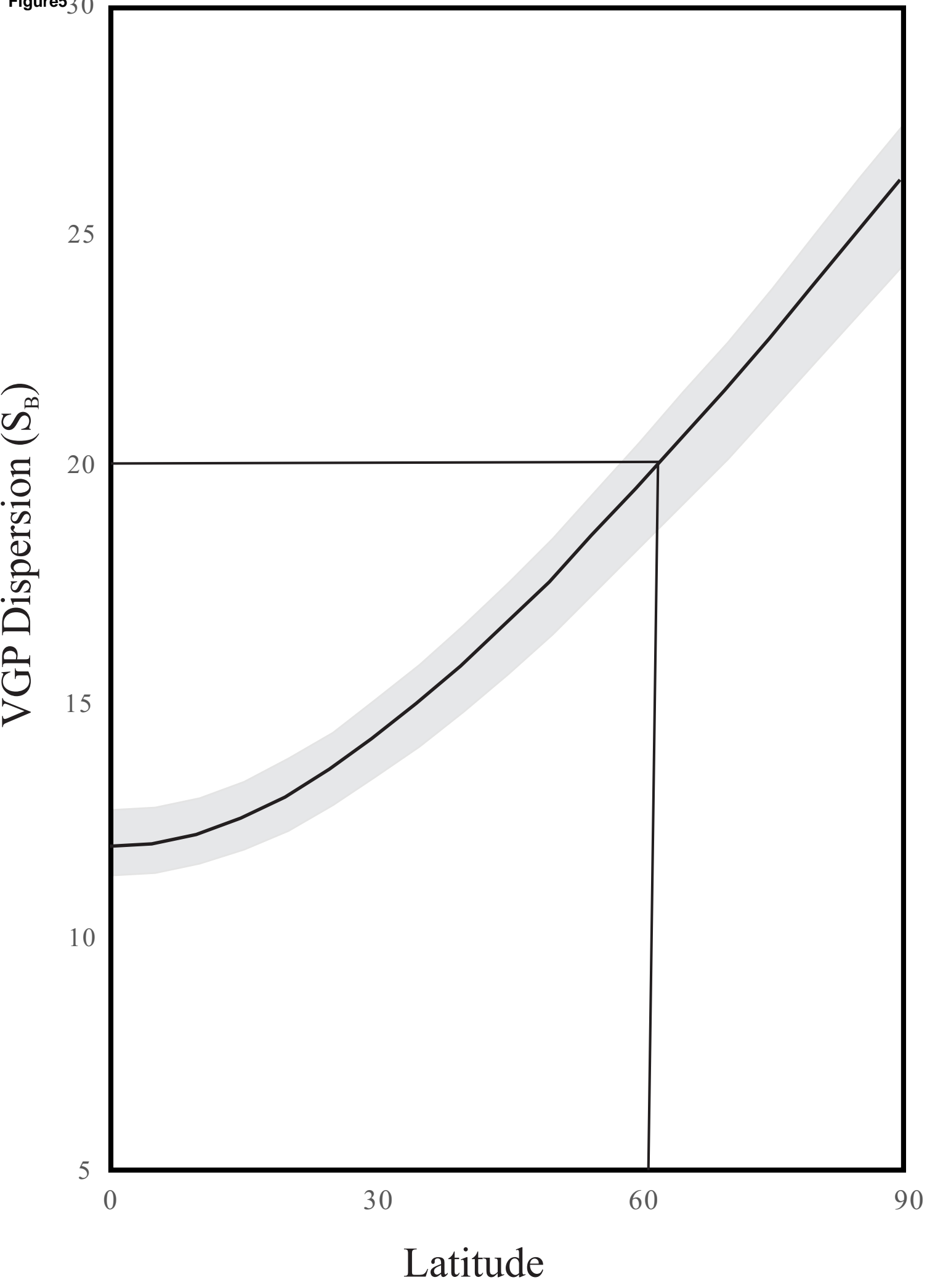
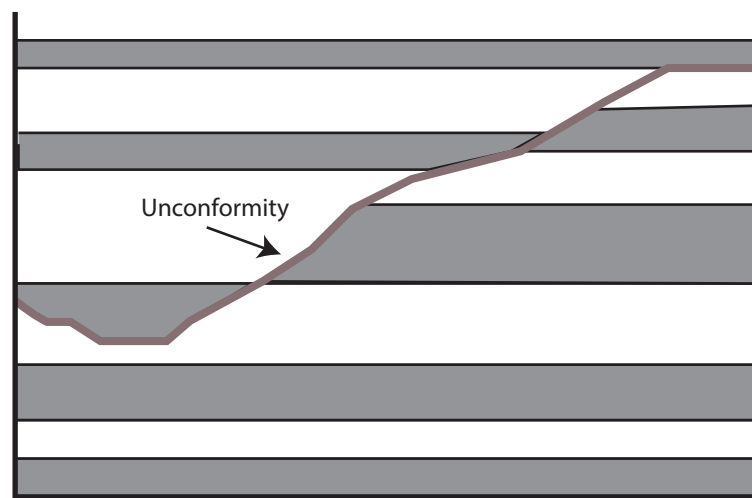
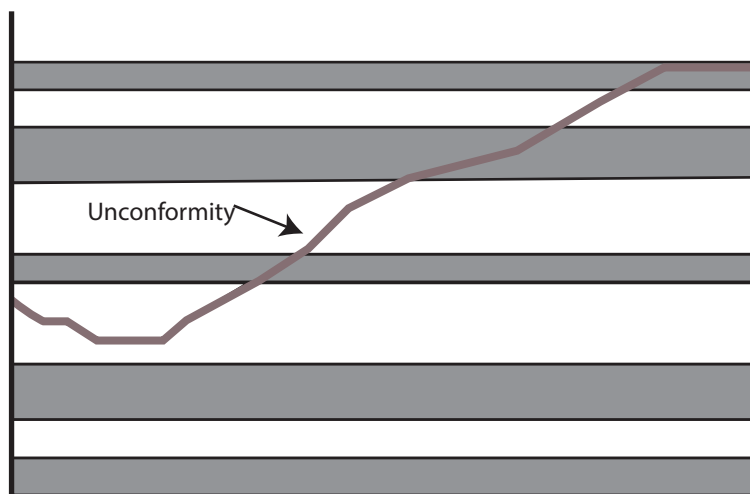
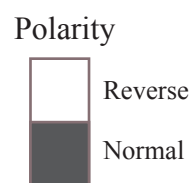


Figure6



(a)



(b)

Figure7

Inclination Flattening

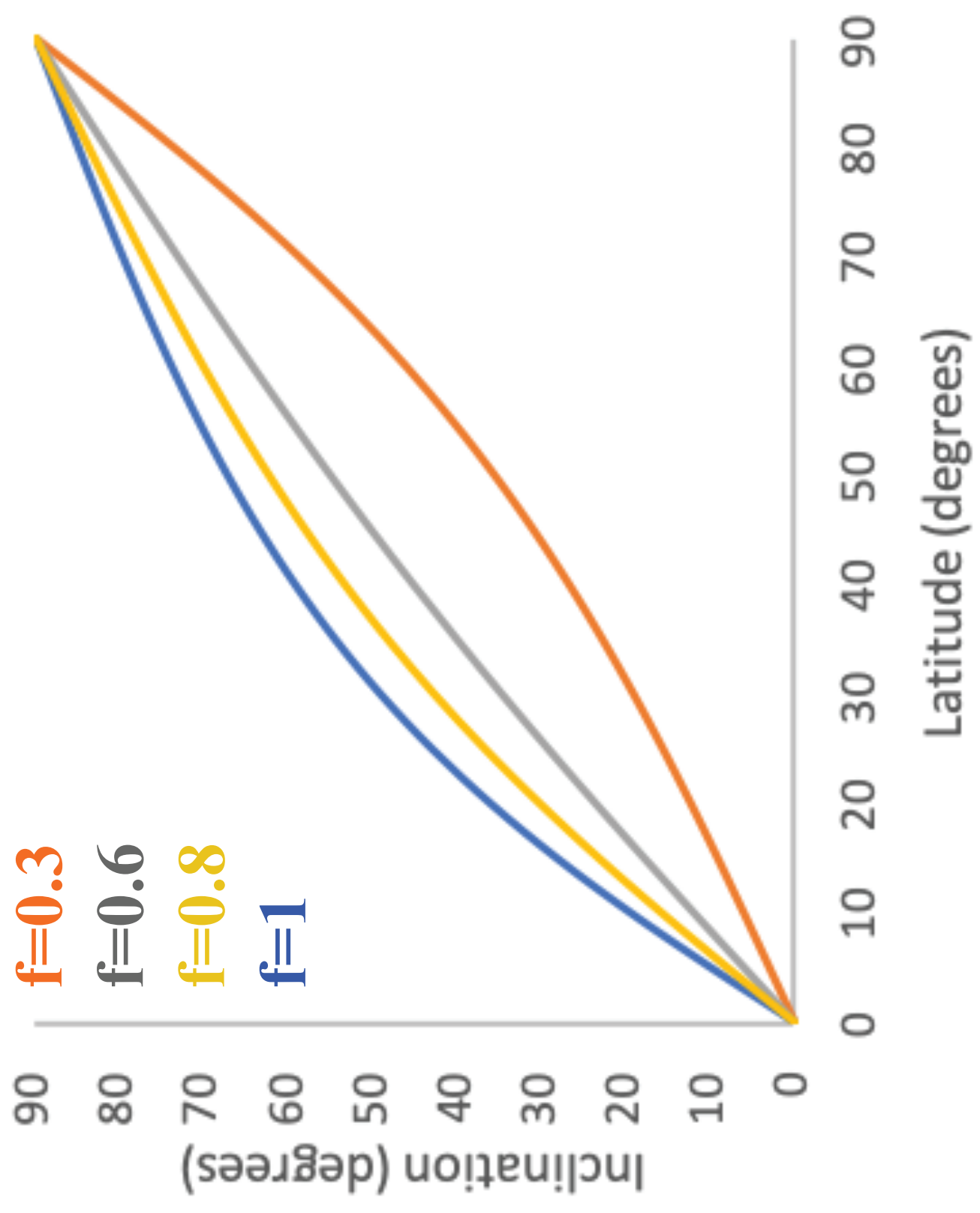


Figure8

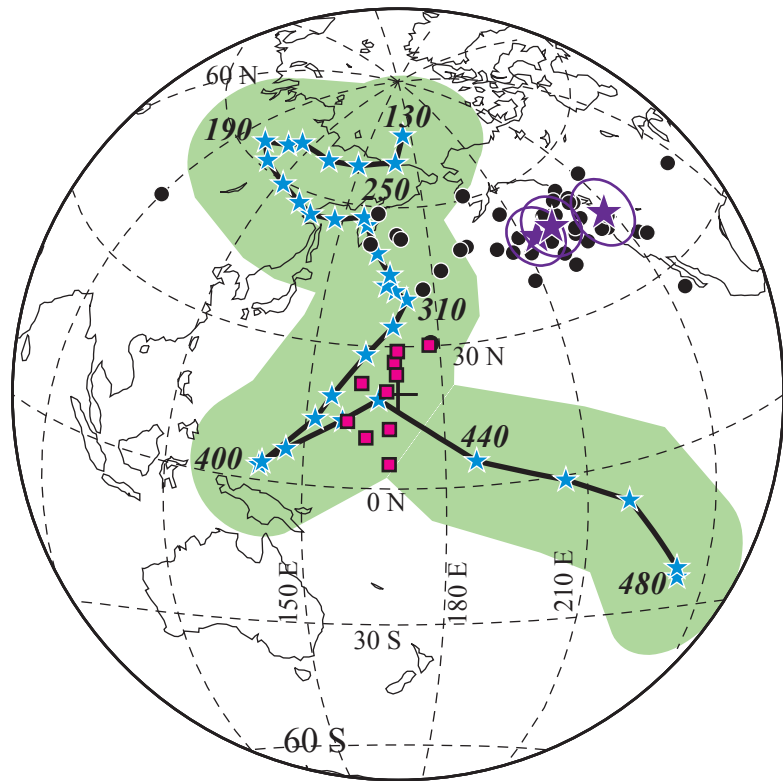
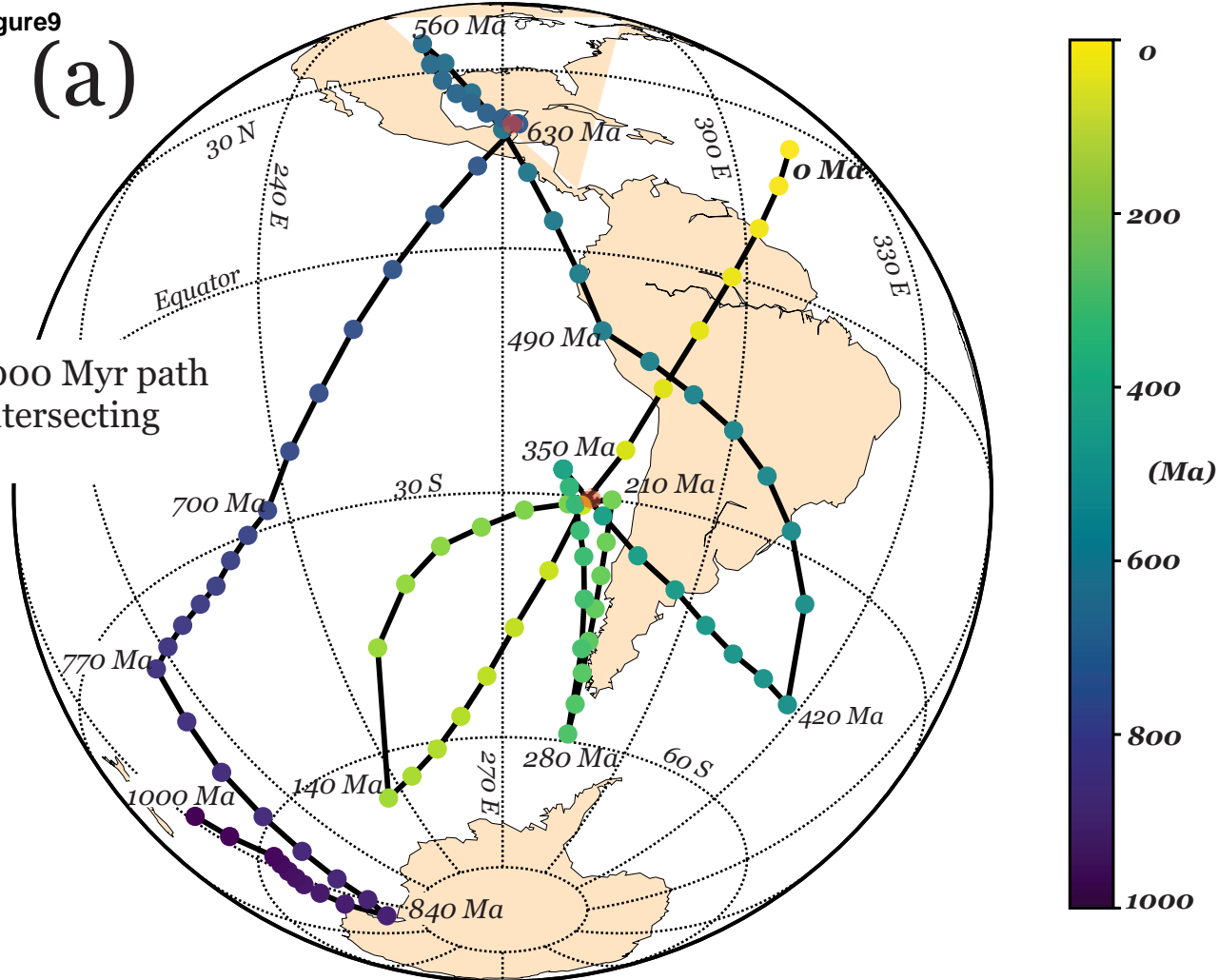


Figure9

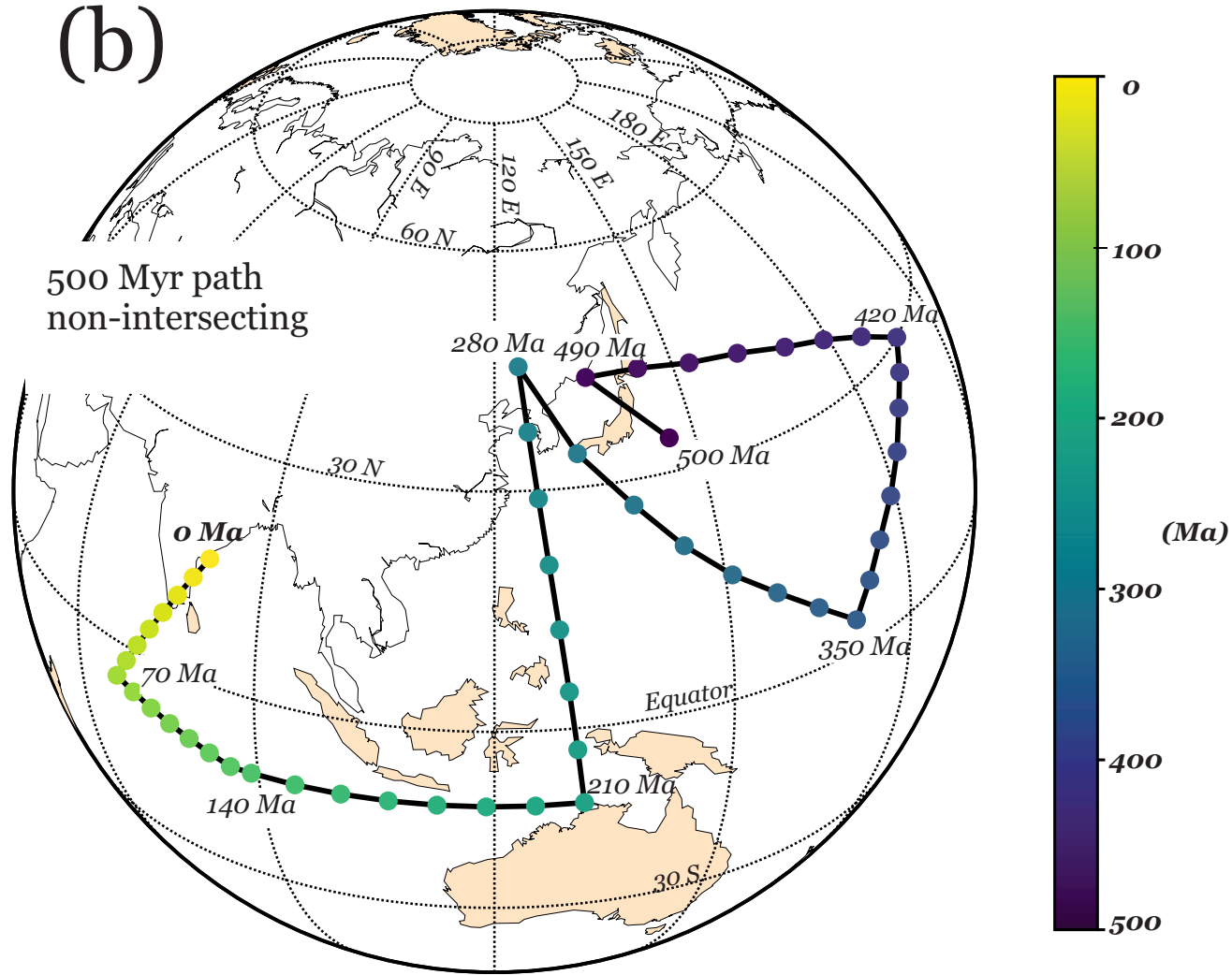
(a)

1000 Myr path intersecting



(b)

500 Myr path non-intersecting



Declaration of interests

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

The authors declare the following financial interests/personal relationships which may be considered as potential competing interests: